

p1

Simplified Solar Sail.

1/31/2022

$$\text{Sun: } L_0 = 3.8 \times 10^{26} \text{ J/s}$$

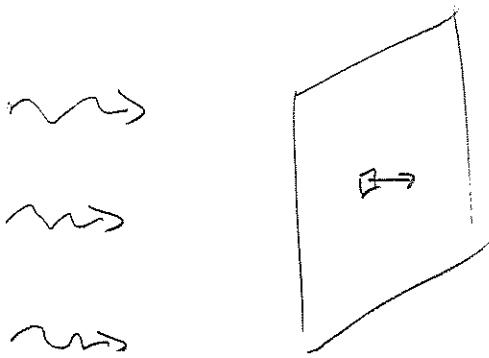
Assume all photons have $\lambda = 5000 \text{ \AA} = 500 \text{ nm}$.

$$\text{Energy per photon } E_\nu = \frac{1240 \text{ eV}\cdot\text{nm}}{500 \text{ nm}} = 2.48 \text{ eV} \times \frac{1.609 \times 10^{-19} \text{ J}}{1 \text{ eV}} = 4 \times 10^{-19} \text{ J}$$

$$\text{Solar flux is } N = \frac{L_0}{E_\nu} = 9.5 \times 10^{44} \frac{\text{photon}}{\text{s}}$$

So at distance R from Sun, flux is

$$f = \frac{N}{4\pi R^2} = \frac{9.5 \times 10^{44} \frac{\text{photon}}{\text{s}}}{4\pi R^2} = \frac{7.58 \times 10^{43} \frac{\text{photon}}{\text{s}\cdot\text{m}^2}}{R^2}$$



Sail of area A is normal to incident solar radiation. Then it intercepts

$$\text{photons intercepted} = \frac{7.58 \times 10^{43} \frac{\text{photon}}{\text{s}\cdot\text{m}^2}}{R^2} \cdot A$$

Each photon is absorbed (no reflection for this sail).

$$\text{Momentum of 1 photon } p_\nu = \frac{E}{c} = \frac{hc/\lambda}{c} = \frac{h}{\lambda} = \frac{6.626 \times 10^{-34} \text{ kg}\cdot\text{m}^2/\text{s}}{500 \times 10^{-9} \text{ m}}$$

$$= 1.33 \times 10^{-27} \text{ kg}\cdot\text{m/s} / \text{photon}$$

Total momentum absorbed per second

$$\begin{aligned} &= \left(1.33 \times 10^{-27} \frac{\text{kg}\cdot\text{m}}{\text{s}} \right) \left(\frac{7.58 \times 10^{43} \frac{\text{photon}}{\text{s}\cdot\text{m}^2}}{R^2} \cdot A \right) \\ &= \frac{1.00 \times 10^{17} \text{ kg}\cdot\text{m/s}}{R^2 \frac{\text{s}\cdot\text{m}^2}} \cdot A \end{aligned}$$

p2 But change in momentum per second is equal to force on sail:

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$$\int F(t) dt = F \int dt = \Delta p$$

$$\Rightarrow \frac{\Delta p}{\Delta t} = F$$

So force due to 100% absorbed photons is

$$F_{\text{abs}} = 1.00 \times 10^{17} \frac{A}{R^2} \text{ Newton}$$

Example: $R = 1 \text{ AU} = 1.5 \times 10^{11} \text{ m}$

$$A = 1 \text{ m} \times 1 \text{ m} = 1 \text{ m}^2$$

$$F_{\text{abs}} = 1.00 \times 10^{17} \frac{1 \text{ m}^2}{(1.5 \times 10^{11} \text{ m})^2} = 4.46 \times 10^{-5} \text{ N}$$

If sail made from aluminum foil, of thickness W and density ρ

$$\rho = 2700 \text{ kg/m}^3$$

$$W = 100 \mu\text{m} = 10^{-4} \text{ m}$$

$$\text{Mass of sail } m = A \cdot W \cdot \rho = 1 \text{ m}^2 \cdot 10^{-4} \text{ m} \cdot 2700 \frac{\text{kg}}{\text{m}^3} = 0.27 \text{ kg}$$

$$\text{Force of grav } F_g = \frac{GM_{\odot} m}{R^2} = \frac{(6.67 \times 10^{-11} \frac{\text{N} \cdot \text{m}^2}{\text{kg}^2})(1.99 \times 10^{30} \text{ kg})(0.27 \text{ kg})}{(1.5 \times 10^{11} \text{ m})^2}$$

$$= 0.0016 \text{ N} = 1.6 \times 10^{-3} \text{ N}$$

In this case, grav > photon pressure

p2

Real sail materials

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IKAROS 7.5 μm polyimide $\sigma \approx 10 \text{ g/m}^2$
 $= 10^{-2} \text{ kg/m}^2$

LightSail 2 4.6 μm Mylar

"Salisbury Screen" $\sigma \approx 3 \text{ g/m}^2$
 $= 3 \times 10^{-3} \text{ kg/m}^2$

Note that all these will yield a sail mass so large that grav force $>$ photon pressure force. So none will sail directly away from Sun. However, they can modify a pre-existing orbit, raising or lowering perihelion/aphelion

So, first step is

- assume some value (realistic) for σ . Say, $\sigma = 5 \text{ g/m}^2$
- assume sail always \perp to radial vector
- assume sail absorbs all photons
- assume sail is placed in circular orbit to start

and follow motion of sail over, say, 10 or 100 years.

Computing the components of the force.

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Let's start with coordinates of the Sun and ship

$$\text{Sun: } (x_0, y_0, z_0)$$

$$\text{ship } (x_s, y_s, z_s)$$

1) compute radial vector \vec{r}

$$\vec{r} = (x_s - x_0, y_s - y_0, z_s - z_0)$$

2) compute magnitude of \vec{r}

$$|\vec{r}| = \sqrt{(x_s - x_0)^2 + (y_s - y_0)^2 + (z_s - z_0)^2}$$

3) compute unit vector \hat{r}

$$\hat{r} = \frac{\vec{r}}{|\vec{r}|} = \left(\frac{x_s - x_0}{|\vec{r}|}, \frac{y_s - y_0}{|\vec{r}|}, \frac{z_s - z_0}{|\vec{r}|} \right)$$

Now, suppose you have calculated the size of the photon force on the sail as F_{photon} . Then the components of force are

$$\vec{F} = F_{\text{photon}} \hat{r}$$

$$= F_{\text{photon}} \cdot \left(\frac{x_s - x_0}{|\vec{r}|}, \frac{y_s - y_0}{|\vec{r}|}, \frac{z_s - z_0}{|\vec{r}|} \right)$$

This calculation is correct only if the sail is face-on to the Sun.