

Figure 1 - Simulated Spectrum of Earth if it were observed from 10pc.
Credit: Edwin Alexani, CfD.

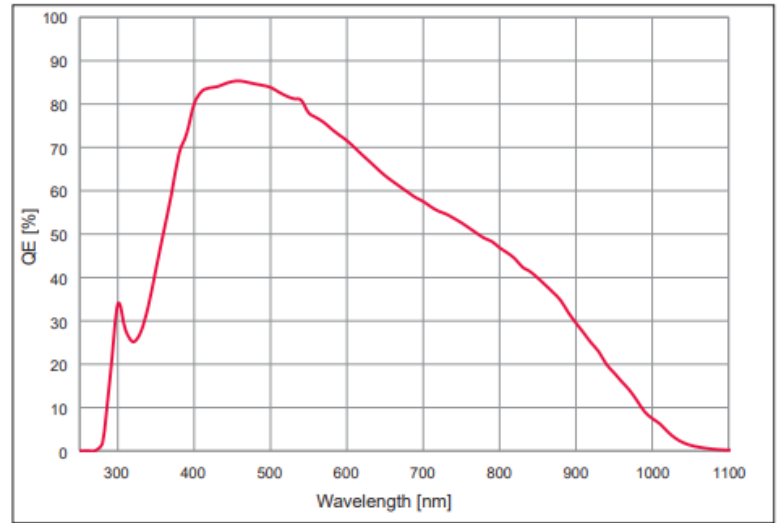


Fig. 2-3. QE curve

Figure 2 - Hamamatsu Orca Quest camera technical note document.²
Hamamatsu.com

Suppose you are planning to do some exoplanet research with an exceptionally low noise CMOS camera such as the ²Hamamatsu QuestTM. Considering the atmosphere of modern earth, a good absorption line to look for is the Oxygen A-band at ~762nm¹. The Hamamatsu Quest has an area of 4096pix x 2304pix. [per channel can be interpreted as per pixel]

- Looking at the simulated spectrum and quantum efficiency seen above, what signal would you be able to measure with an exposure of 1 second [electrons]?
- ²Given a dark current of 0.006 electrons/s/channel and a readout noise of 0.27 electrons_{rms}/channel, estimate the SNR of your measurement.
- To know if there is an actual absorption line at this wavelength, we need a baseline. Estimate the measured signal and SNR you would expect to see at 800nm.

- [Molecular Oxygen – A band absorption](#)
- [Hamamatsu Quest Technical Note](#)
- Hint: Info on SNR can be found in the [lecture notes](#)

- d. Given this difference in signal and an estimate of the accuracy of these measurements, would you be able to definitively claim there is a presence of O₂ in the simulated exoplanet?

- e. Suppose we wanted a higher signal. If the effective temperature of this exoplanet is 255K, what is the peak blackbody wavelength emitted?

- f. Would you be able to make observations in this range with the Hamamatsu Quest™?

1. [Molecular Oxygen – A band absorption](#)
2. [Hamamatsu Quest Technical Note](#)
3. Hint: Info on SNR can be found in the [lecture notes](#)

Solutions

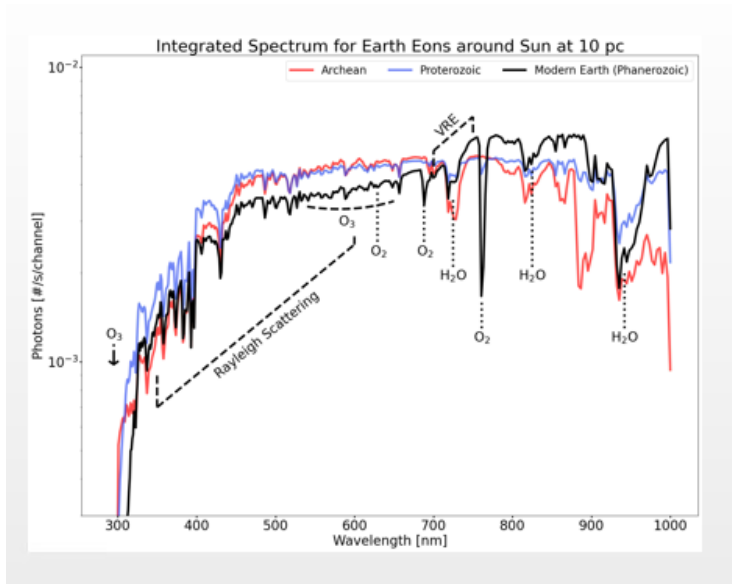


Figure 3 - Simulated Spectrum of Earth if it were observed from 10pc.
Credit: Edwin Alexani, CfD.

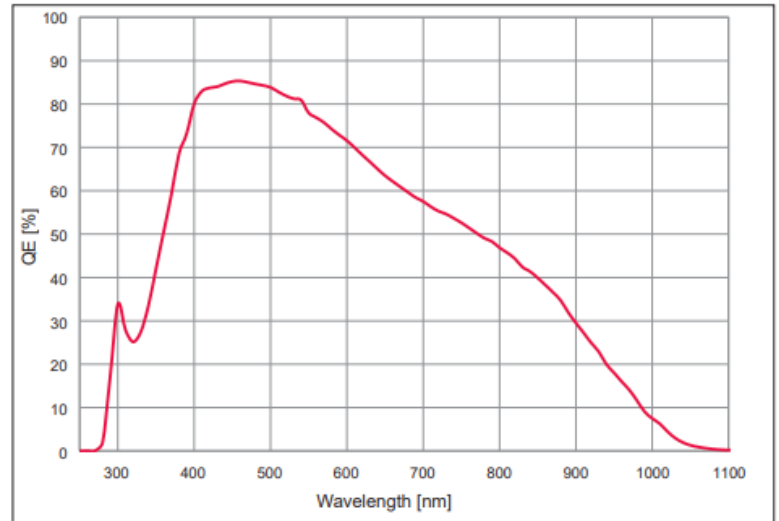


Fig. 2-3. QE curve

Figure 4 - Hamamatsu Orca Quest camera technical note document.
Hamamatsu.com

Suppose you are planning to do some exoplanet research with an exceptionally low noise CMOS camera such as the *Hamamatsu Quest*TM. Considering the atmosphere of modern earth, a good absorption line to look for is the Oxygen A-band at ~762nm¹. The Hamamatsu Quest has an area of 4096pix x 2304pix.

- Looking at the simulated spectrum and quantum efficiency seen above, what signal would you be able to measure with an exposure of 1 second [electrons]?

Incoming photons: $.5 \times 10^{-3}$ photons/s/channel

Integration: $1 \text{ s} * \text{QE}$

Quantum Efficiency: 50%, $0.5 \text{ e}^-/\text{photon}$

$n_{\text{pix}} = 4096 * 2304 = 9437184$ pixels (channels)

Measured Signal: $\text{photons} * \text{integration} * \text{QE} + n_{\text{pix}} = 0.5 \times 10^{-3} * 1 * 0.5 * 9.4 \times 10^6 = 2359 \text{ e}^-$

- Given a dark current of 0.006 electrons/s/channel and a readout noise of 0.27 electrons_{rms}, estimate the SNR of your measurement. (assume no sky signal)

$$\text{SNR} = \frac{\text{Signal}}{\sqrt{\text{Signal} + (\text{Dark Current} * n_{\text{pix}} * t) + \text{Read Noise}^2 * n_{\text{pix}}}}$$

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Solutions

$$= \frac{2359}{\sqrt{2359 + (0.006 * 9.4 \times 10^6 * 1s) + 0.27^2 * 9.4 \times 10^6}} = 2.73 \text{ (yikes!)}$$

- c. To know if there is an actual absorption line at this wavelength, we need a baseline. Estimate the measured signal and SNR you would expect to see at 800nm.

$$\text{Measured Signal: photons} * \text{integration} * \text{QE} + n_{\text{pix}} = 5.0 \times 10^{-3} * 1 * 0.475 * 9.4 \times 10^6 = 22413 \text{ e}^-$$

$$\begin{aligned} \text{SNR} &= \frac{\text{Signal}}{\sqrt{\text{Signal} + (\text{Dark Current} * n_{\text{pix}} * t) + \text{Read Noise}^2 * n_{\text{pix}}}} \\ &= \frac{22413}{\sqrt{22413 + (0.006 * 9.4 \times 10^6 * 1s) + 0.27^2 * 9.4 \times 10^6}} = 25.592 \text{ (better!)} \end{aligned}$$

- d. Given this difference in signal and an estimate of the accuracy of these measurements, would you be able to definitively claim there is a presence of O₂ in the simulated exoplanet?

I would say yes, though the SNR leaves much to be desired, the difference in measured signal is not bad, and we would be able to definitively define a dip in signal at that wavelength. I would say with these approximations, you would not be able to definitively make statements about the abundances, but it would be clearly present in some amount.

- e. Suppose we wanted a higher signal. If the effective temperature of this exoplanet is 255K, what is the peak blackbody wavelength emitted?

$$\text{Wien's Law: } \lambda_{\text{max}} = 2898 [\mu\text{m} * \text{K}] / T_{\text{eff}} [\text{K}] = 2898 / 255 \sim 11 \mu\text{m}$$

- f. Would you be able to make observations in this range with the Hamamatsu Quest™?

No... it's a silicone detector, where the band gap is about 1.1eV or 1μm

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3. Hint: Info on SNR can be found in the [lecture notes](#)